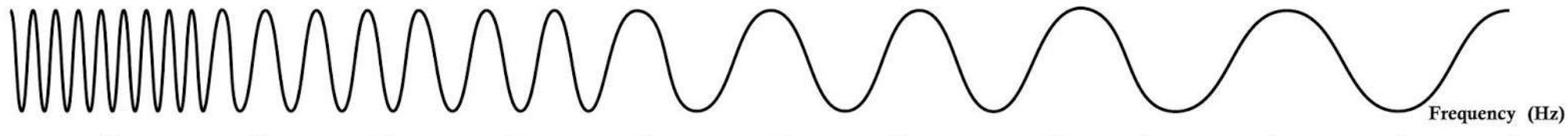


Introduction to Radio Astronomy and it's Effects on our Daily Living...

What does Radio Astronomy
have over Optical?



$10^{24} \quad 10^{22} \quad 10^{20} \quad 10^{18} \quad 10^{16}$ Visible light $10^{14} \quad 10^{12} \quad 10^{10} \quad 10^8 \quad 10^6 \quad 10^4 \quad 10^2$

Gamma rays

X-rays

Uv rays

Infrared rays

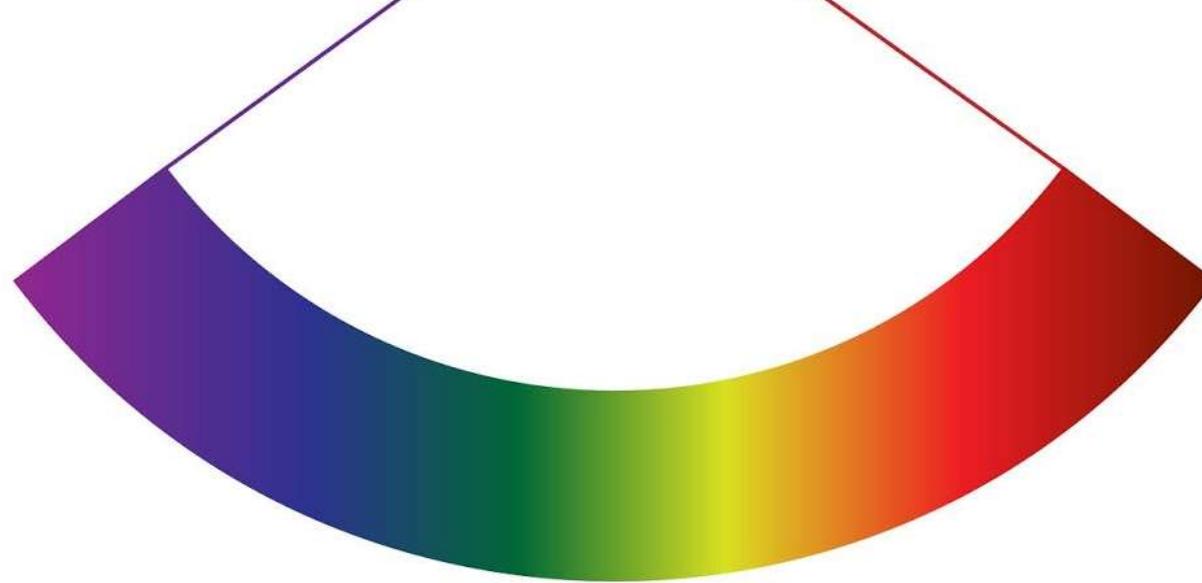
Microwave

Radio, TV

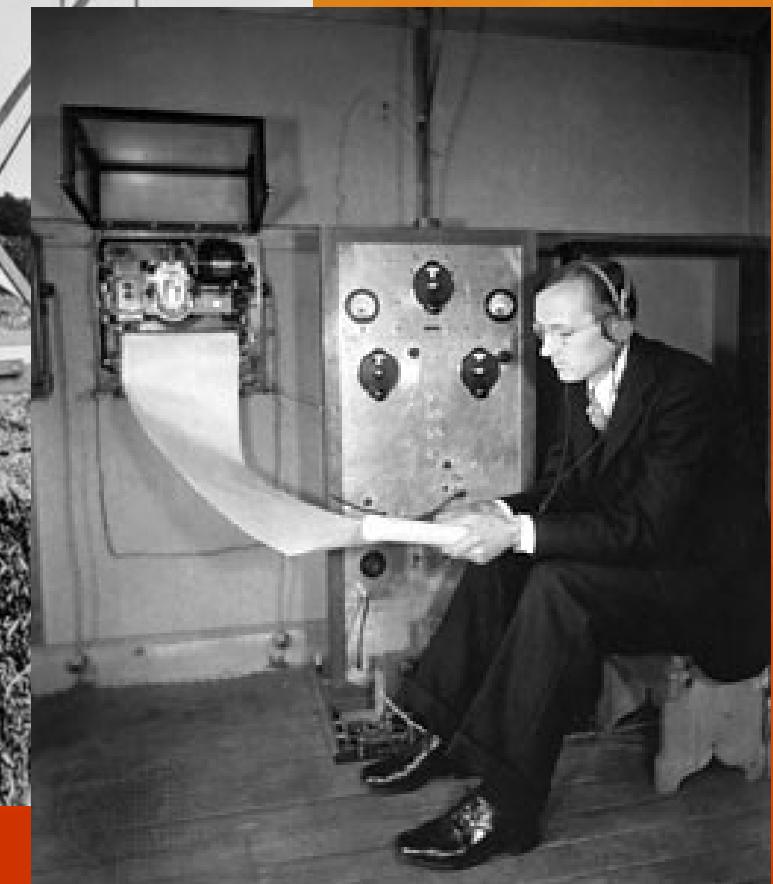
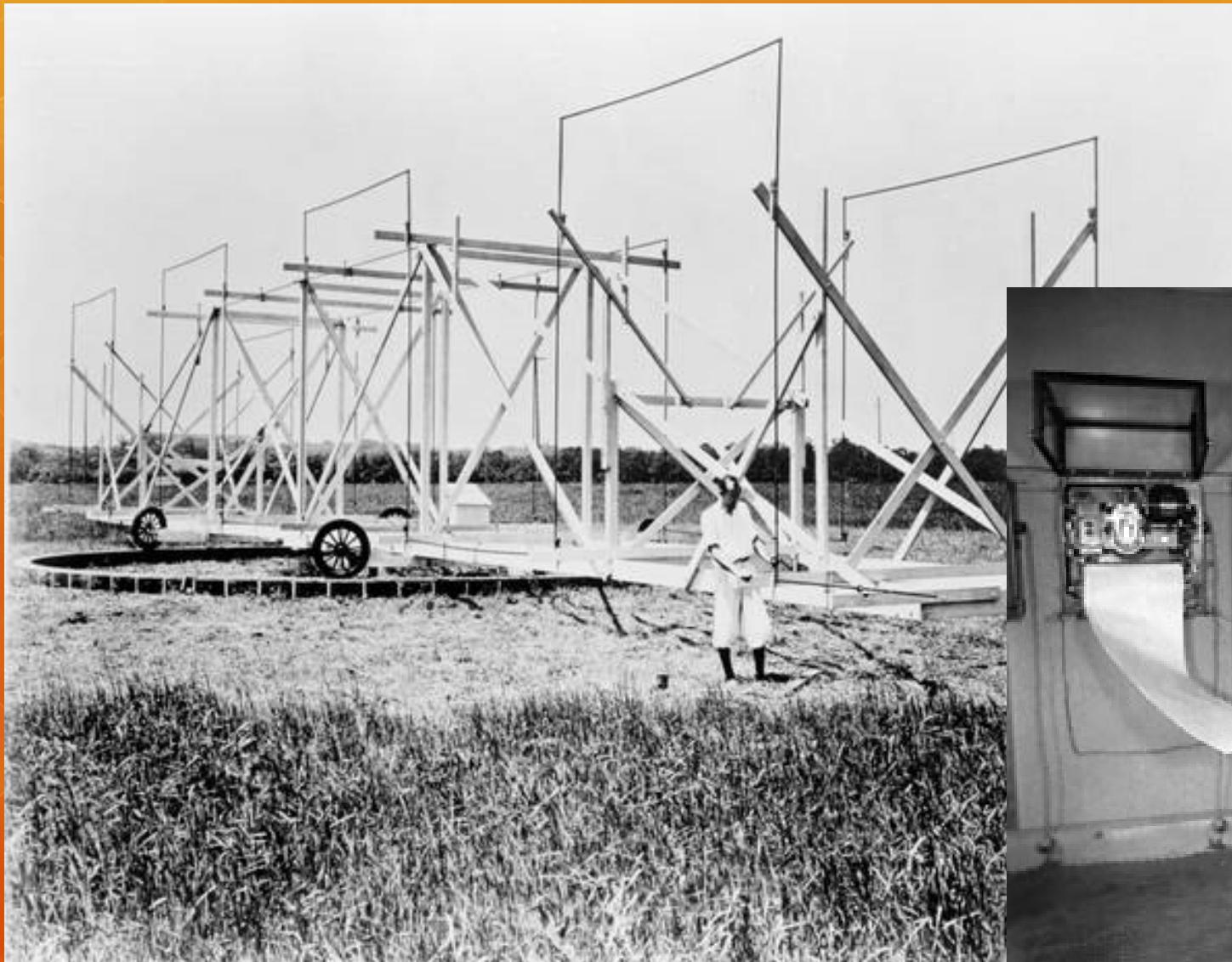
Long radio waves

$10^{-16} \quad 10^{-14} \quad 10^{-12} \quad 10^{-10} \quad 10^{-8} \quad 10^{-6}$ $10^{-4} \quad 10^{-2} \quad 10^0 \quad 10^2 \quad 10^4 \quad 10^6$

Wavelength (m)



Father of Radio Astronomy



Grote Reber



First Radio Graph of our Galaxy

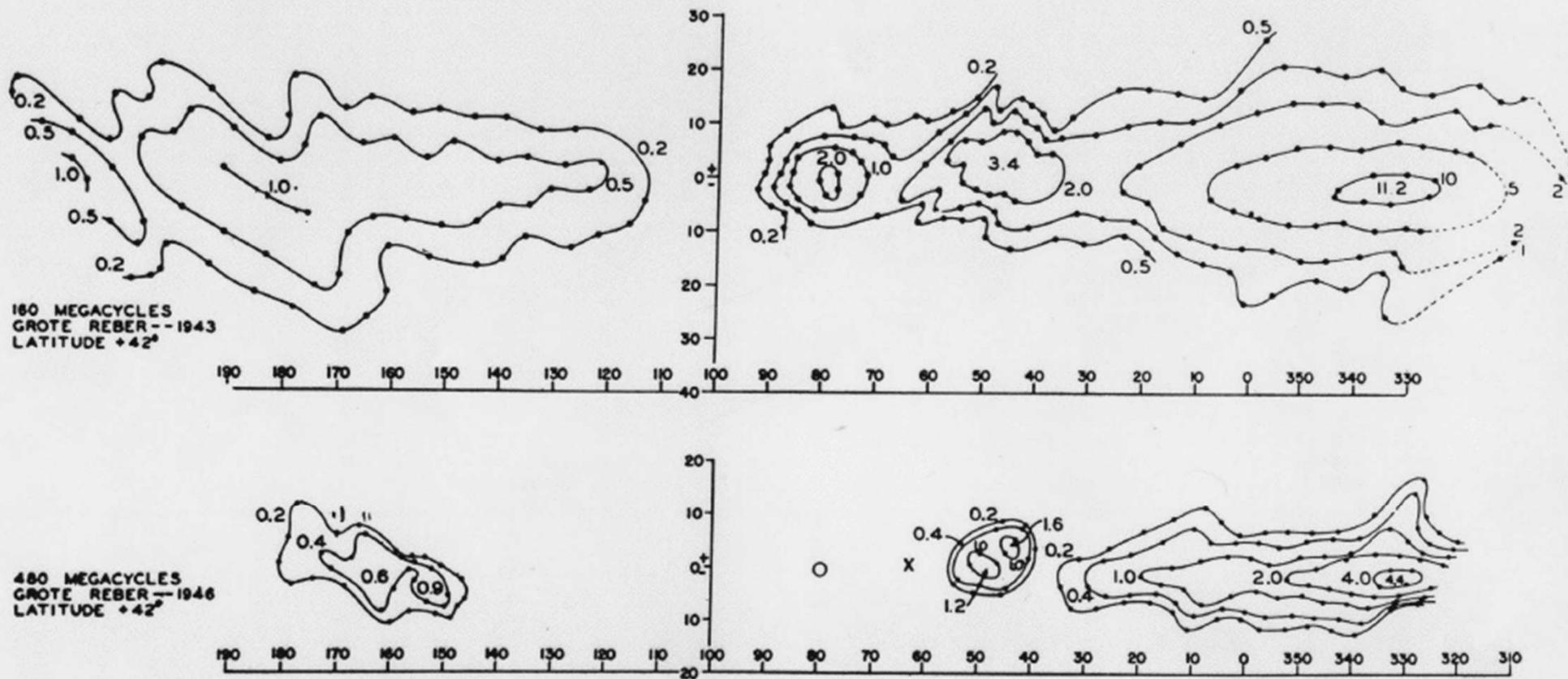
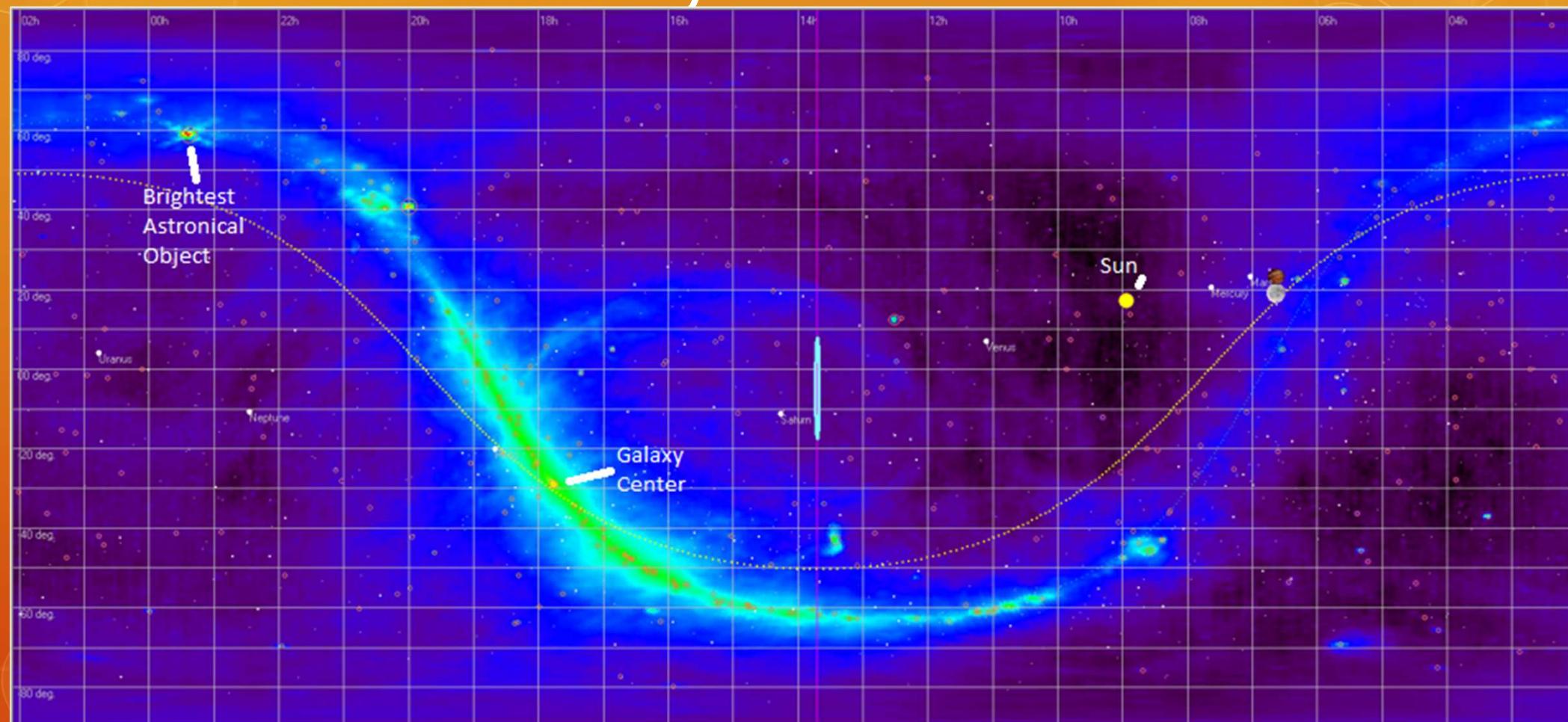


FIG. 7—Contours of constant intensity at 160 MHz and 480 MHz, taken at Wheaton, Illinois.

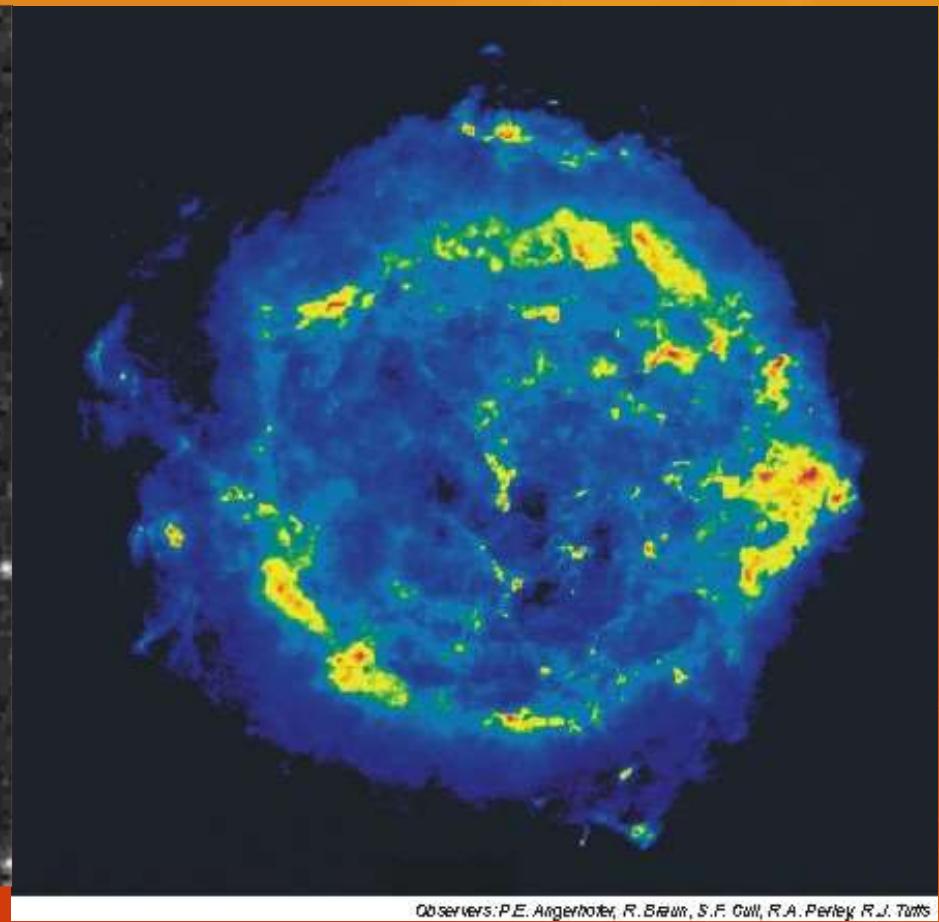
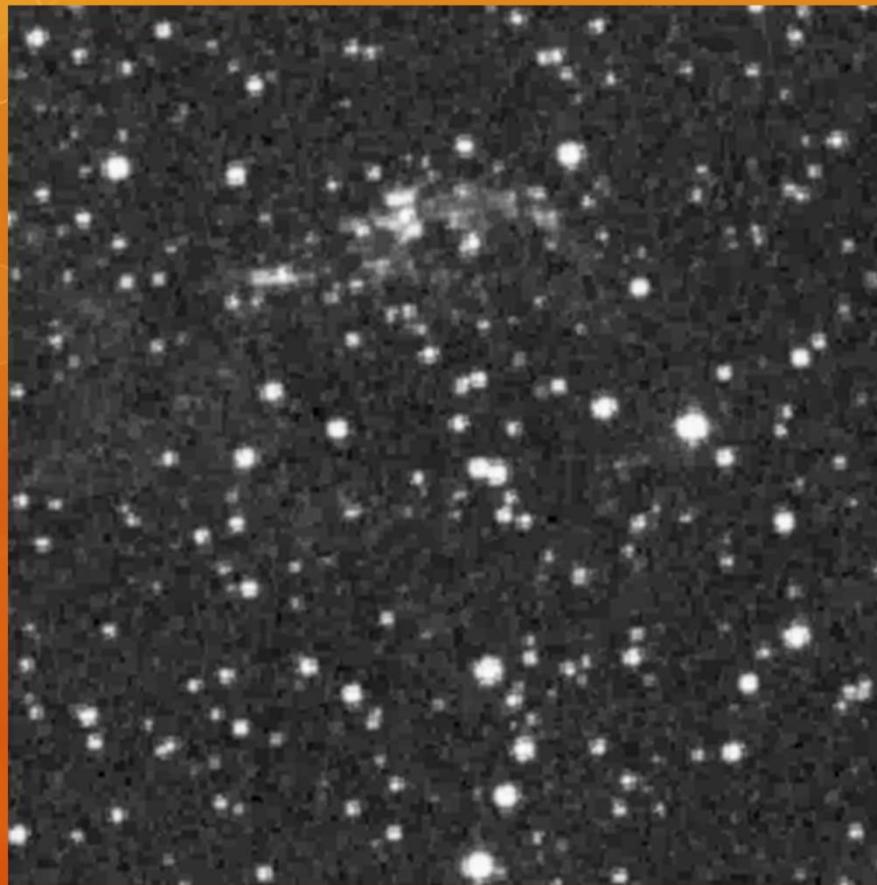
State of the art at the time in the 1930's...

- Astronomers KNEW the universe was thermal.
- Radio was some new-fangled technology that was not going to see much due to the low emission as predicted by the Stefan-Boltzmann equation.
- Radio telescopes have poor resolution. (Initial scopes had resolution measured in degrees)
- Most professional astronomers knew nothing about radio, so they assumed whatever was being received was probably just local noise.

The Whole Sky at 408 MHz.



Supernova Remnant Cassiopeia A

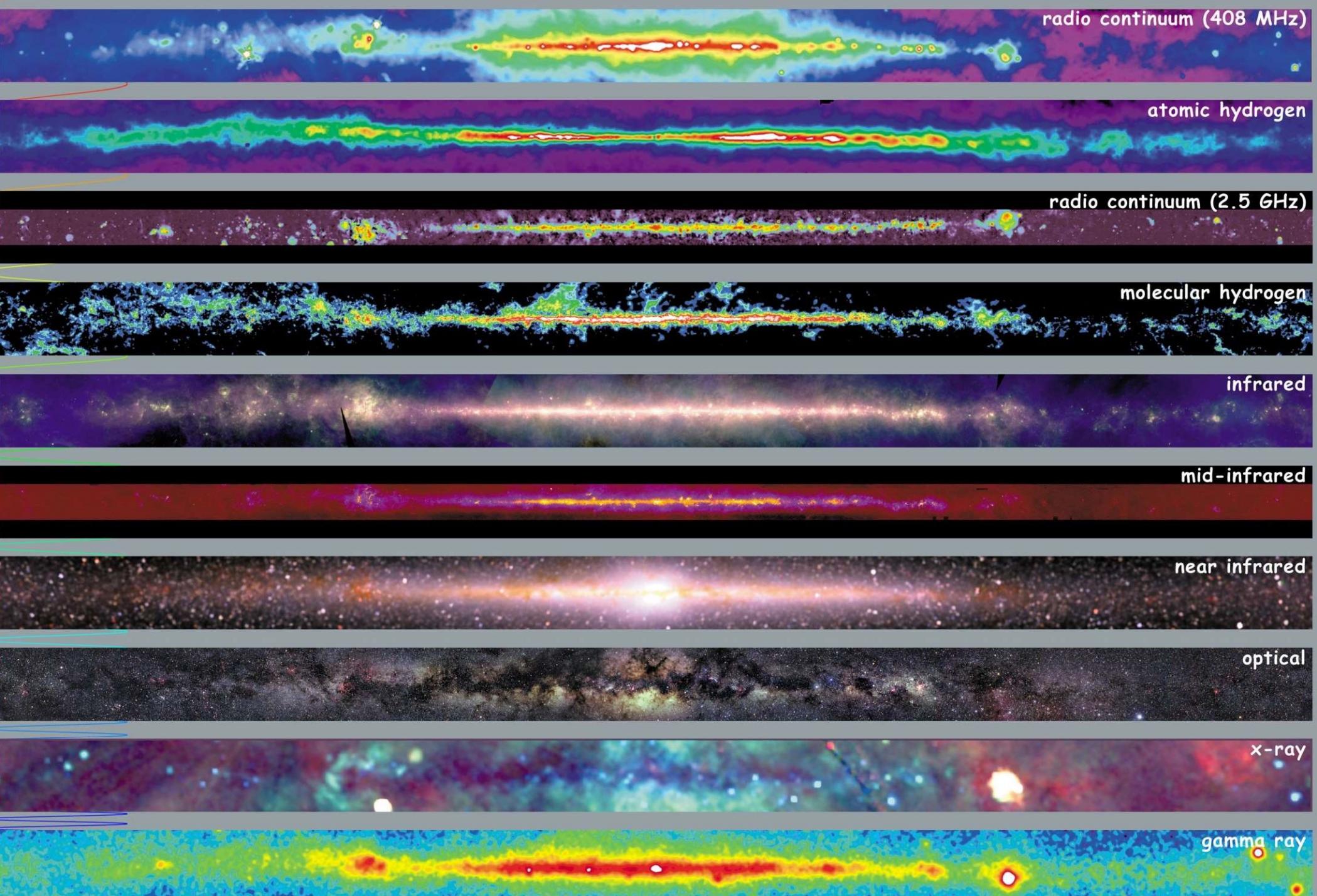


Observers: P.E. Arge, R. Braun, S.F. Gull, R.A. Perley, R.J. Tufts

Radio Galaxy Cygnus A

Cyg A
-Optical/Radio-





Multiwavelength Milky Way

What the professionals use

Image courtesy of NRAO/AU



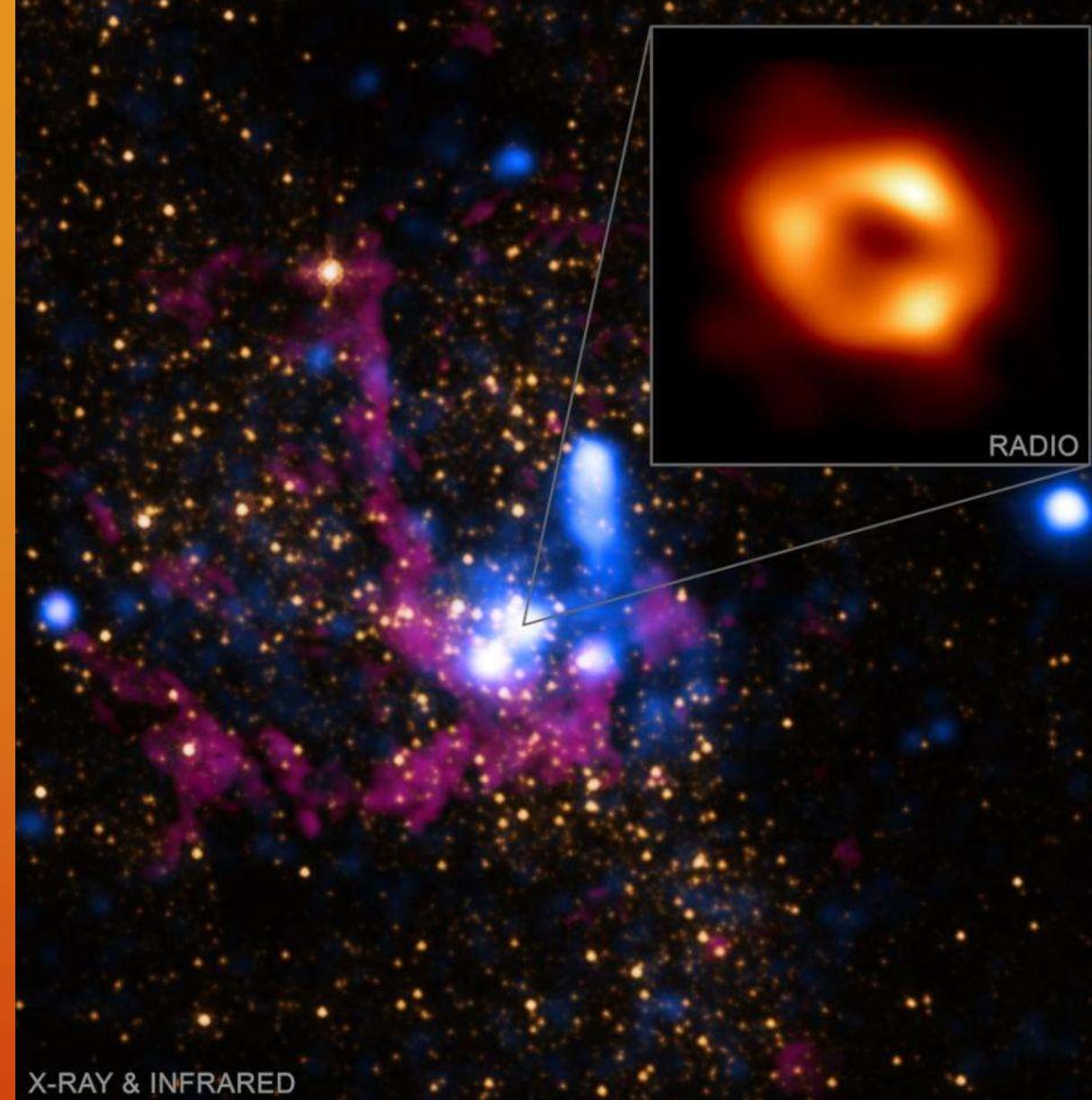
VLA



Event Horizon Telescope



M82



Sagittarius A*

What do the amateurs use?

Eyepieces...

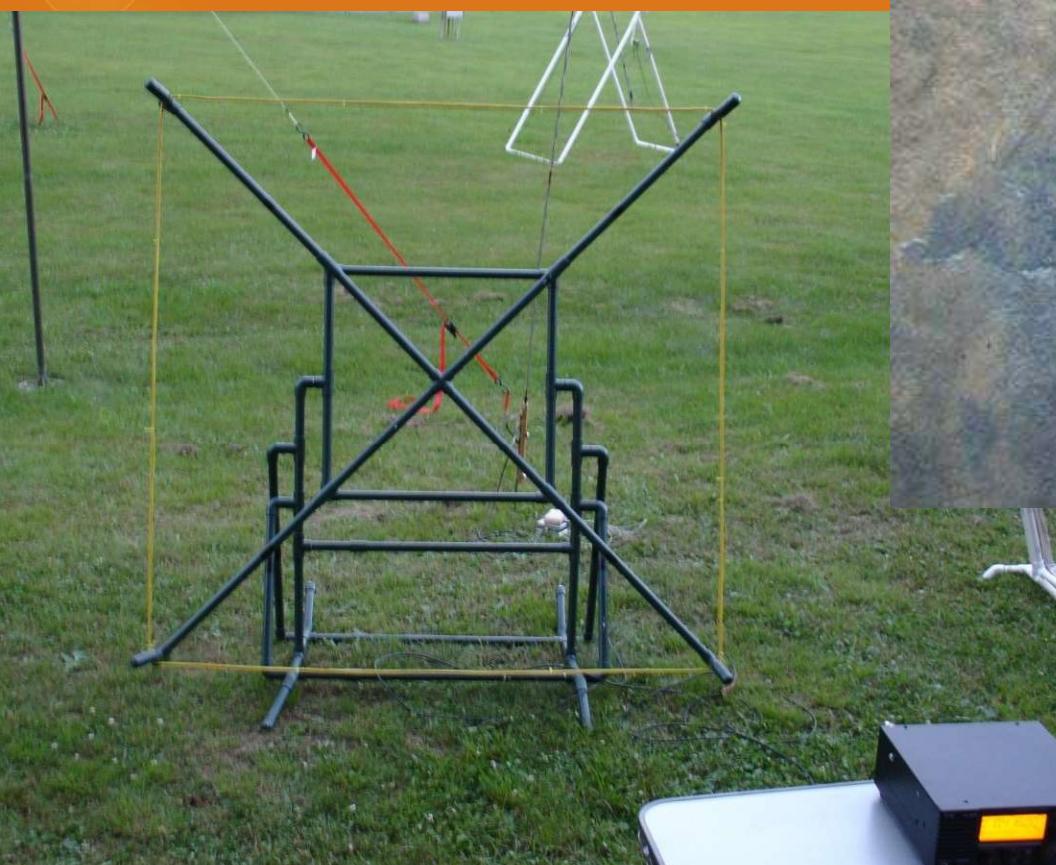


Radio Jove

<http://radiojove.gsfc.nasa.gov/>

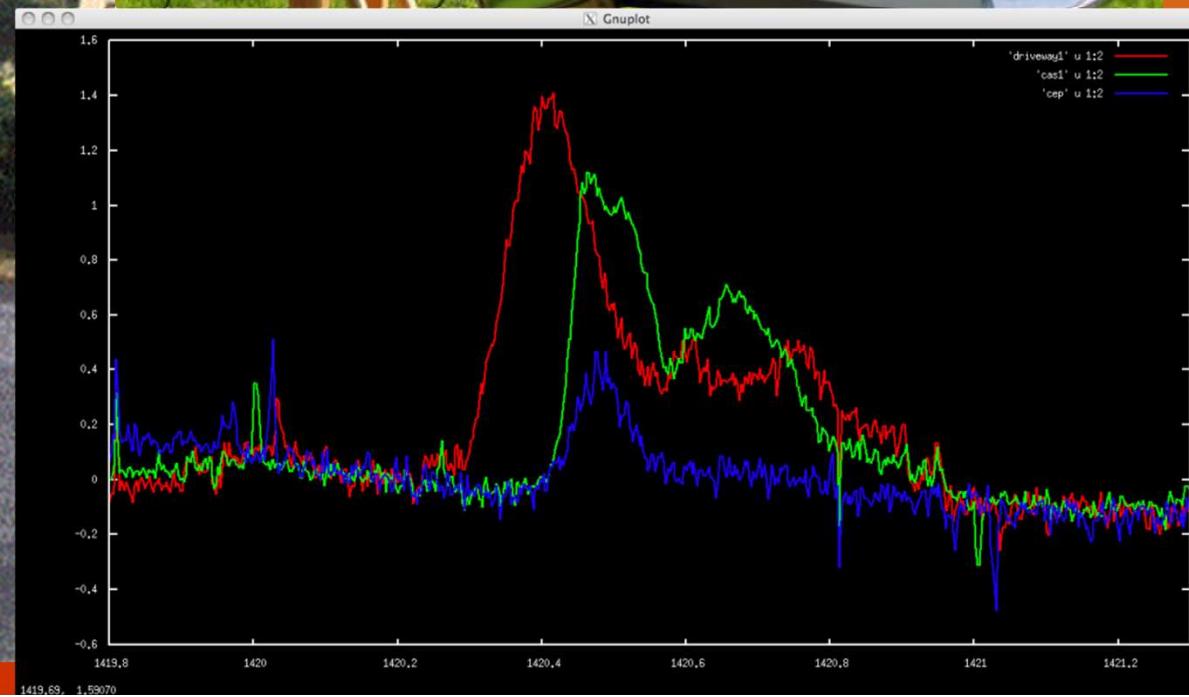


SuperSID

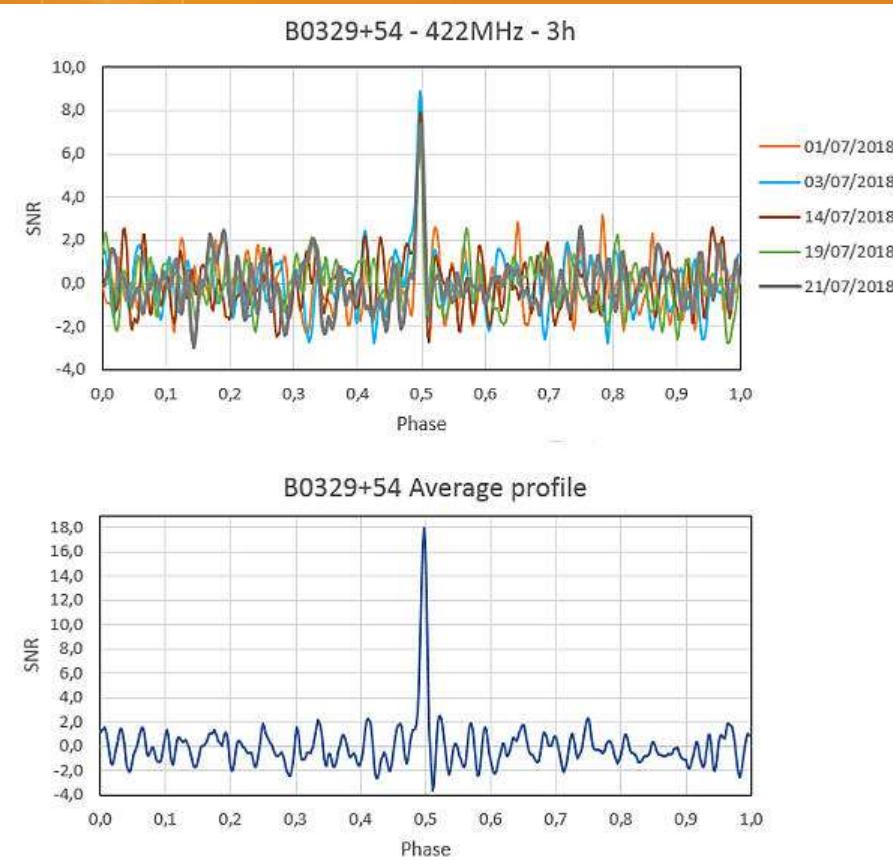


Hydrogen Line

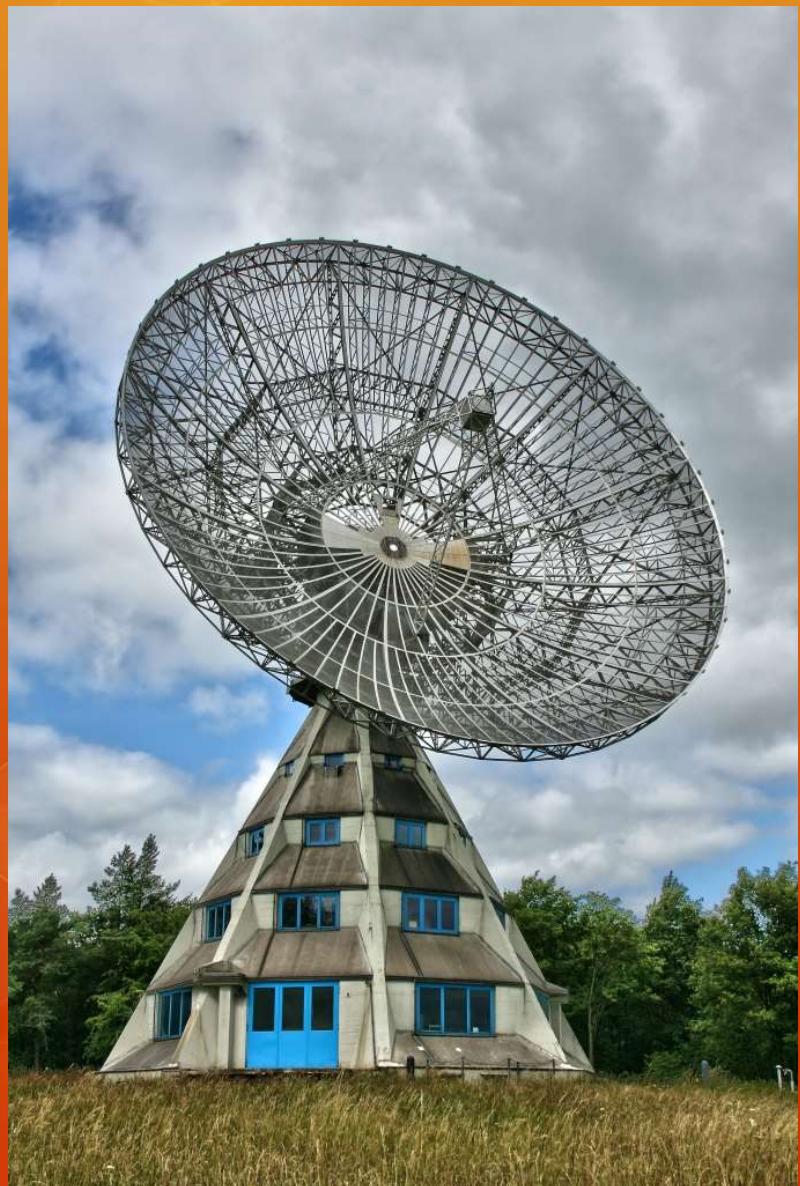
<http://wvurail.org>



Pulsar detection is possible.

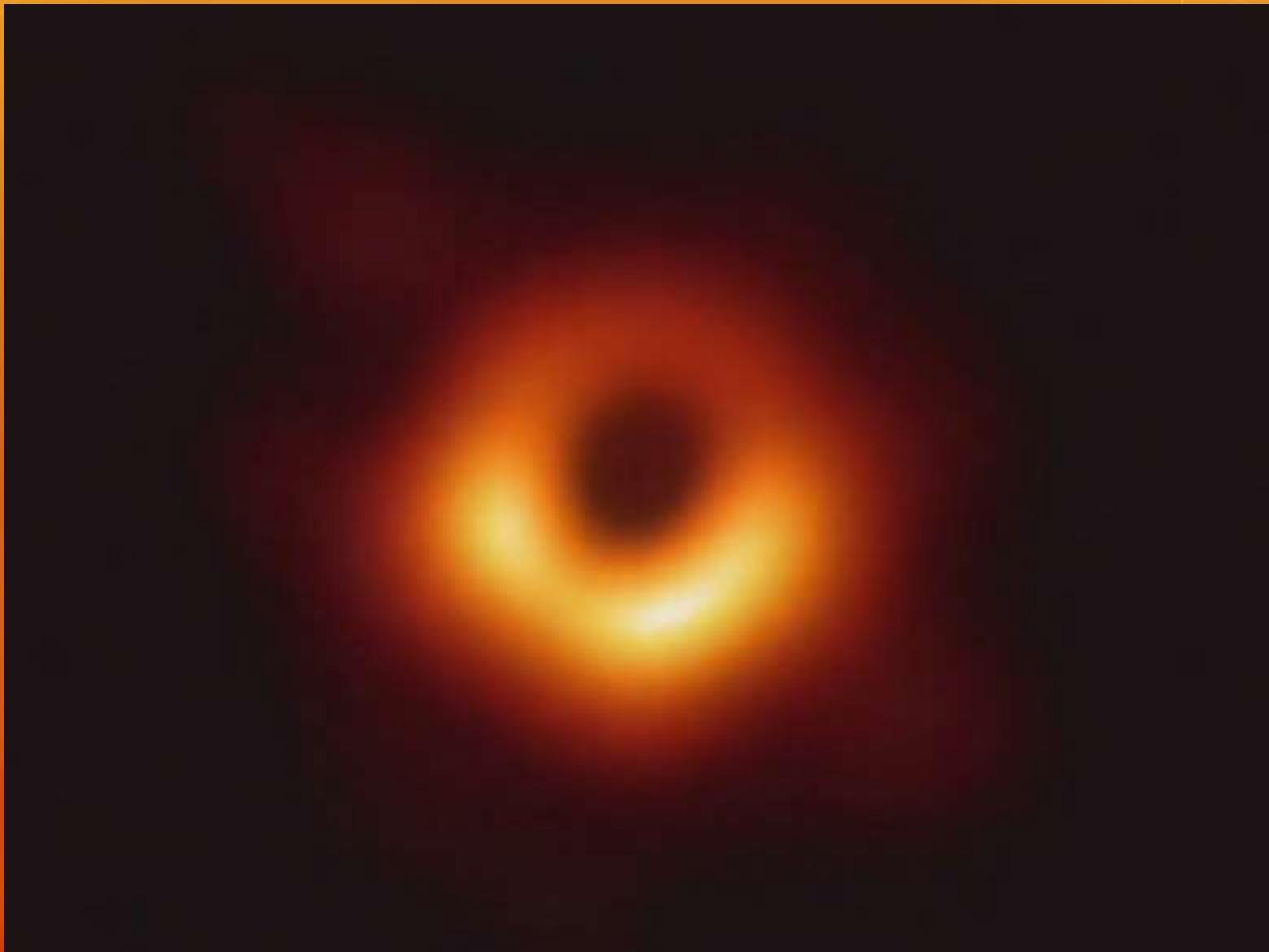


Bigger toys...



What has radio astronomy added
to our knowledge of our
universe?

Imaging Black Holes

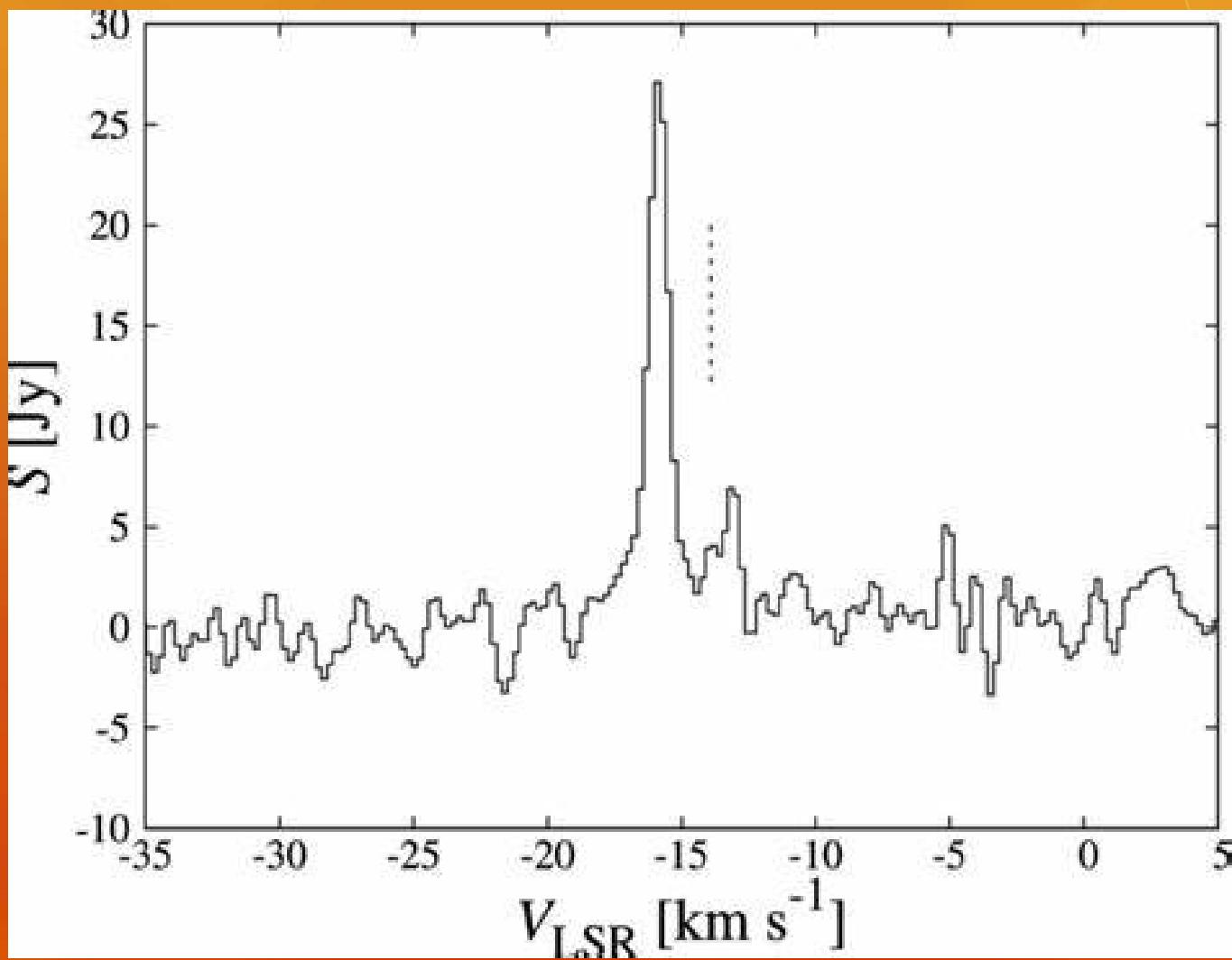


Non-Thermal Radiation from Galaxies

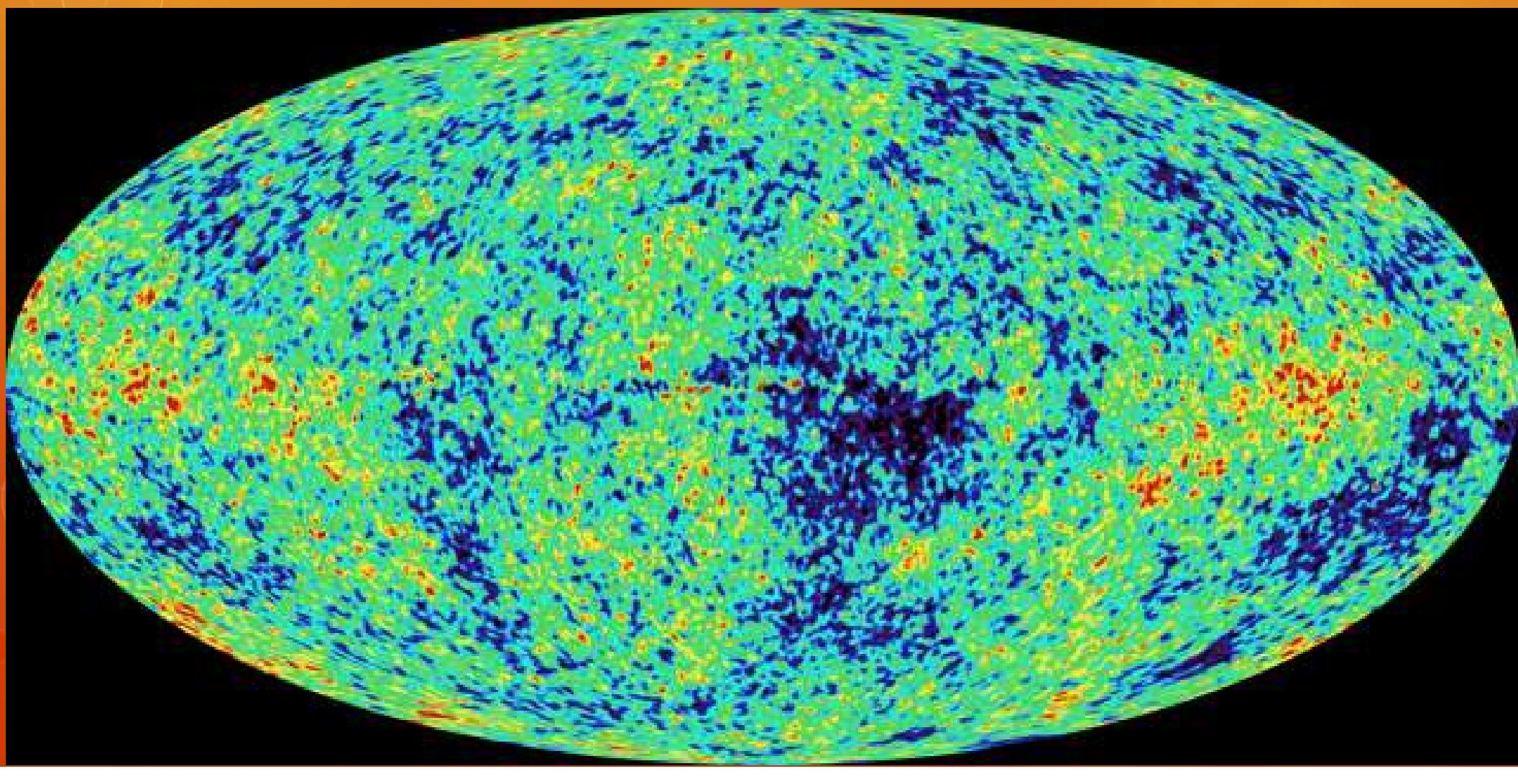


Cyg A
-Optical/Radio-

Coherent Maser Line Emission



Cosmic Microwave Background



Neutron Stars (aka Pulsars)

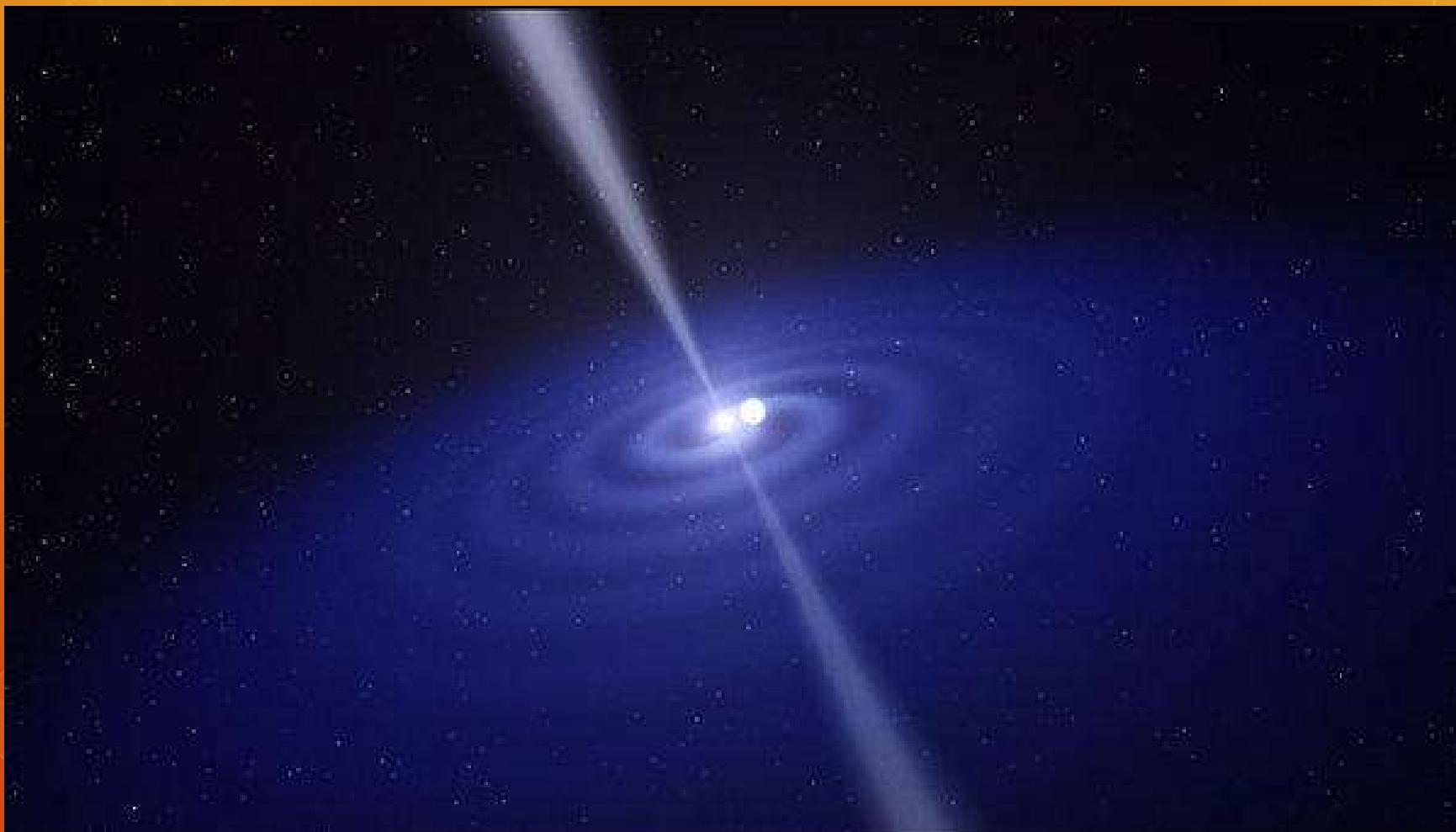
B1933+16
0.356 S



B0531+21
33.4 mS



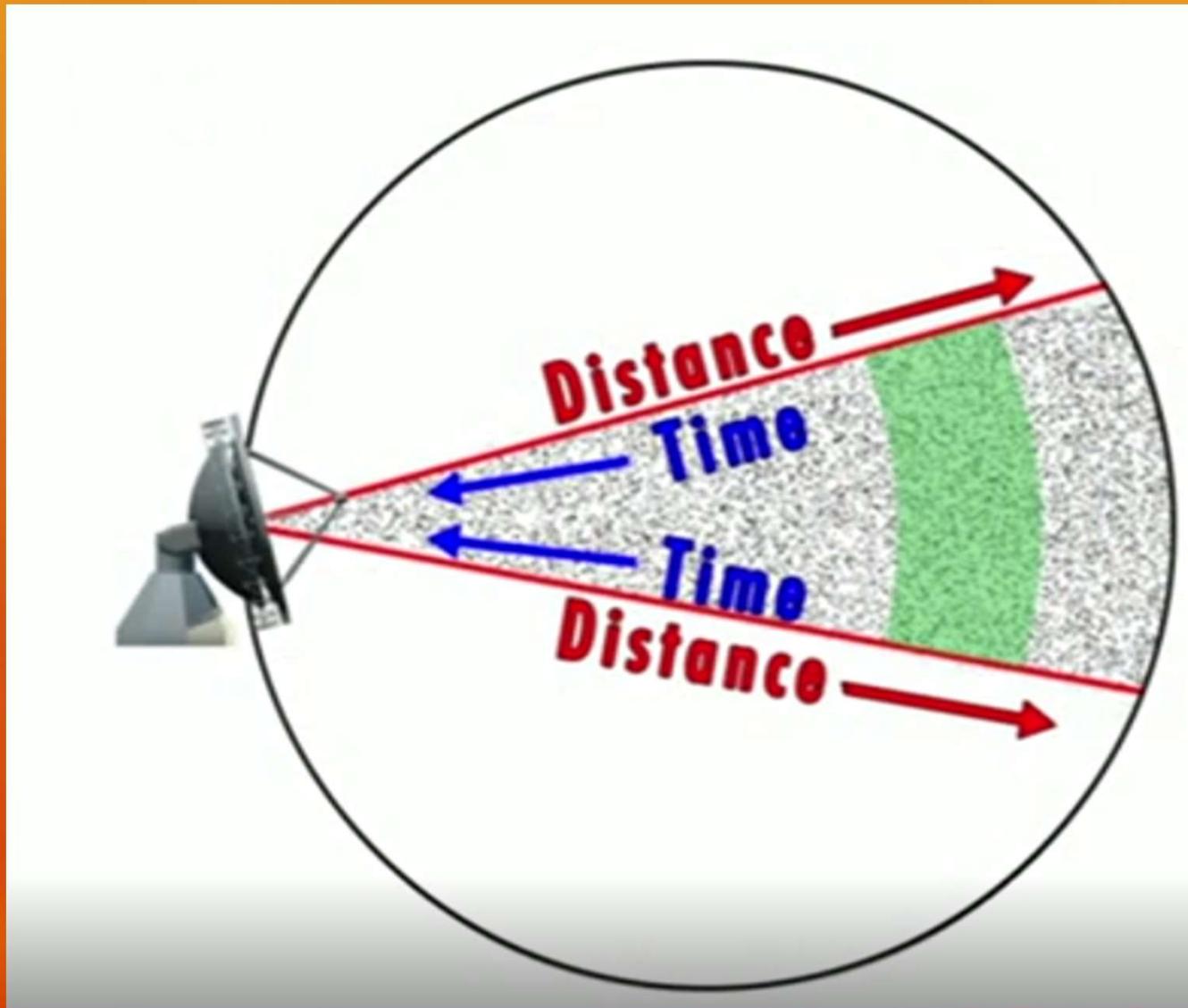
B1937+21
1.56mS



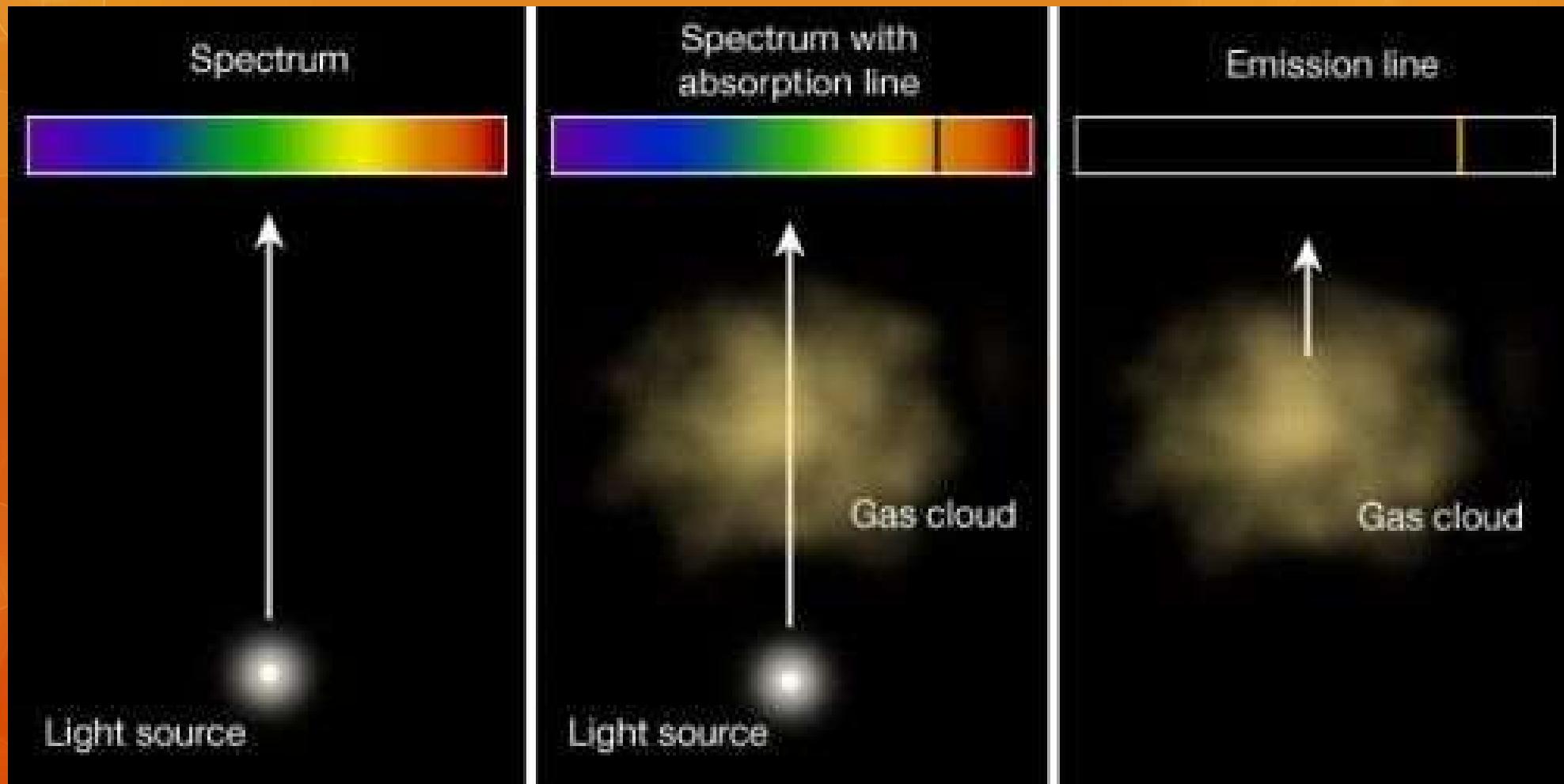
Galaxies extend way beyond their visible area...



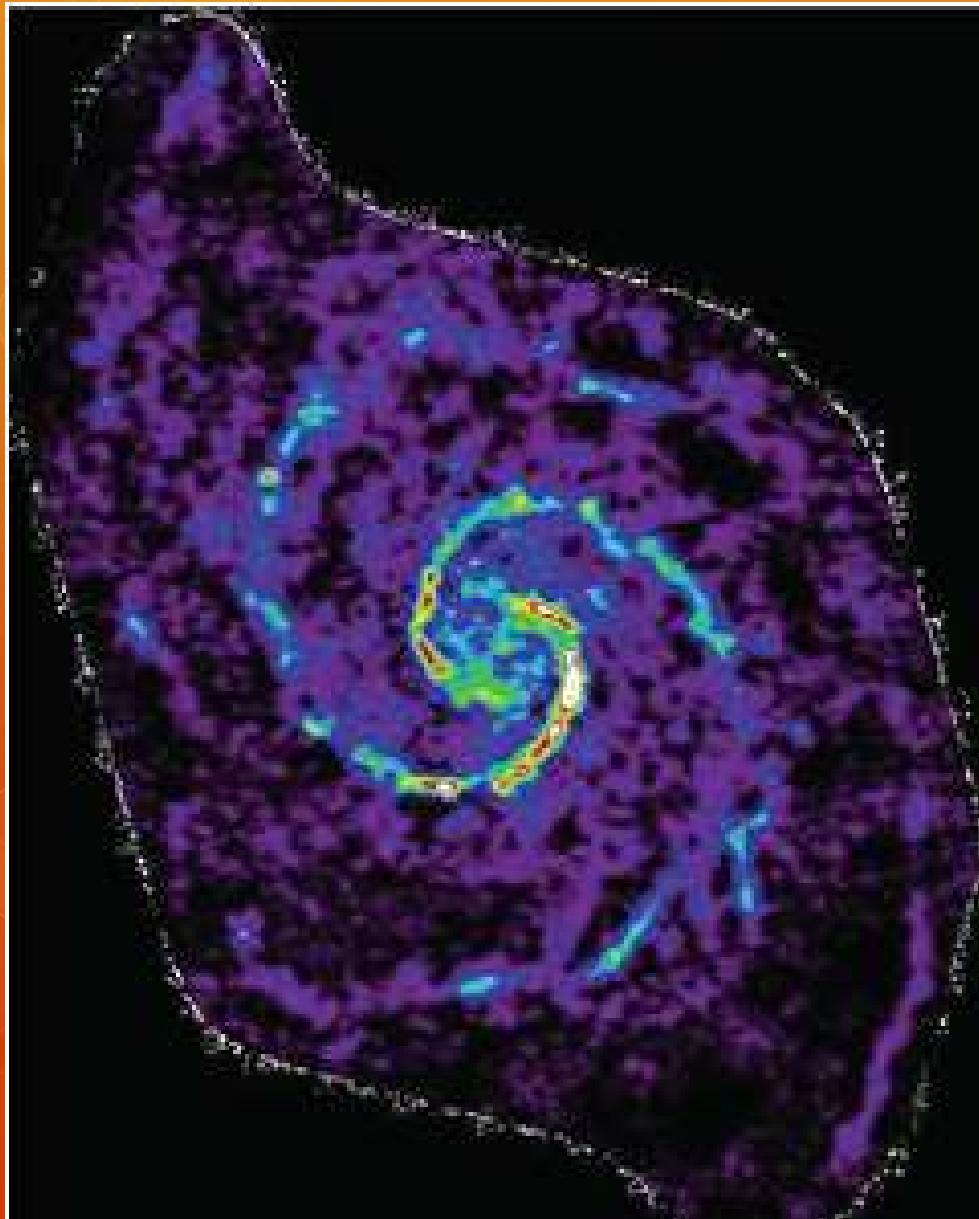
Evolution of Radio Galaxies and Quasars



Thermal spectral-line absorption and emission



Cold areas of galaxies



A peek into the nursery...

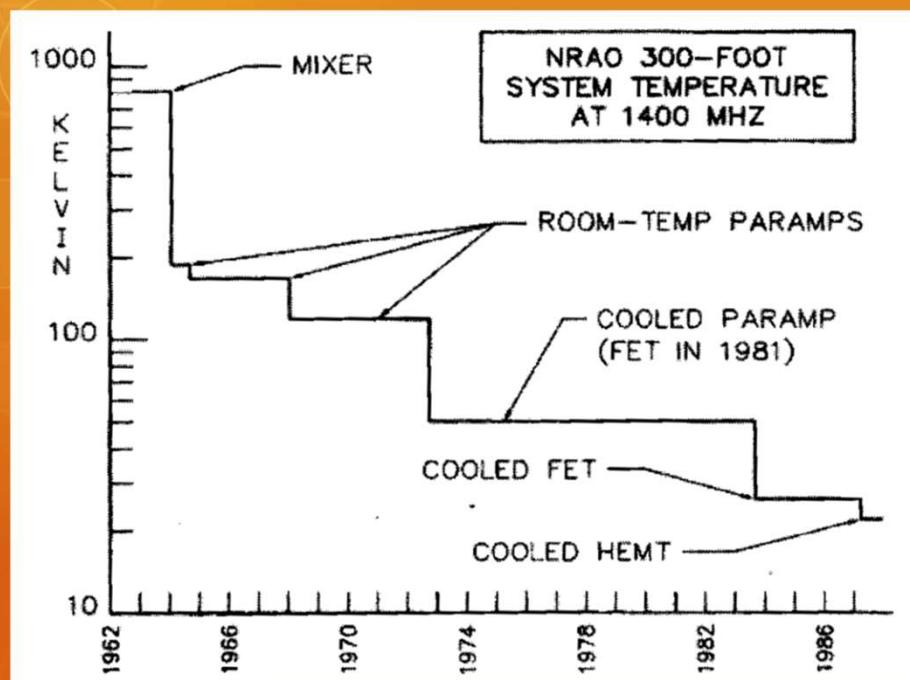


Technologies started in
Radio Astronomy in
everyday use...

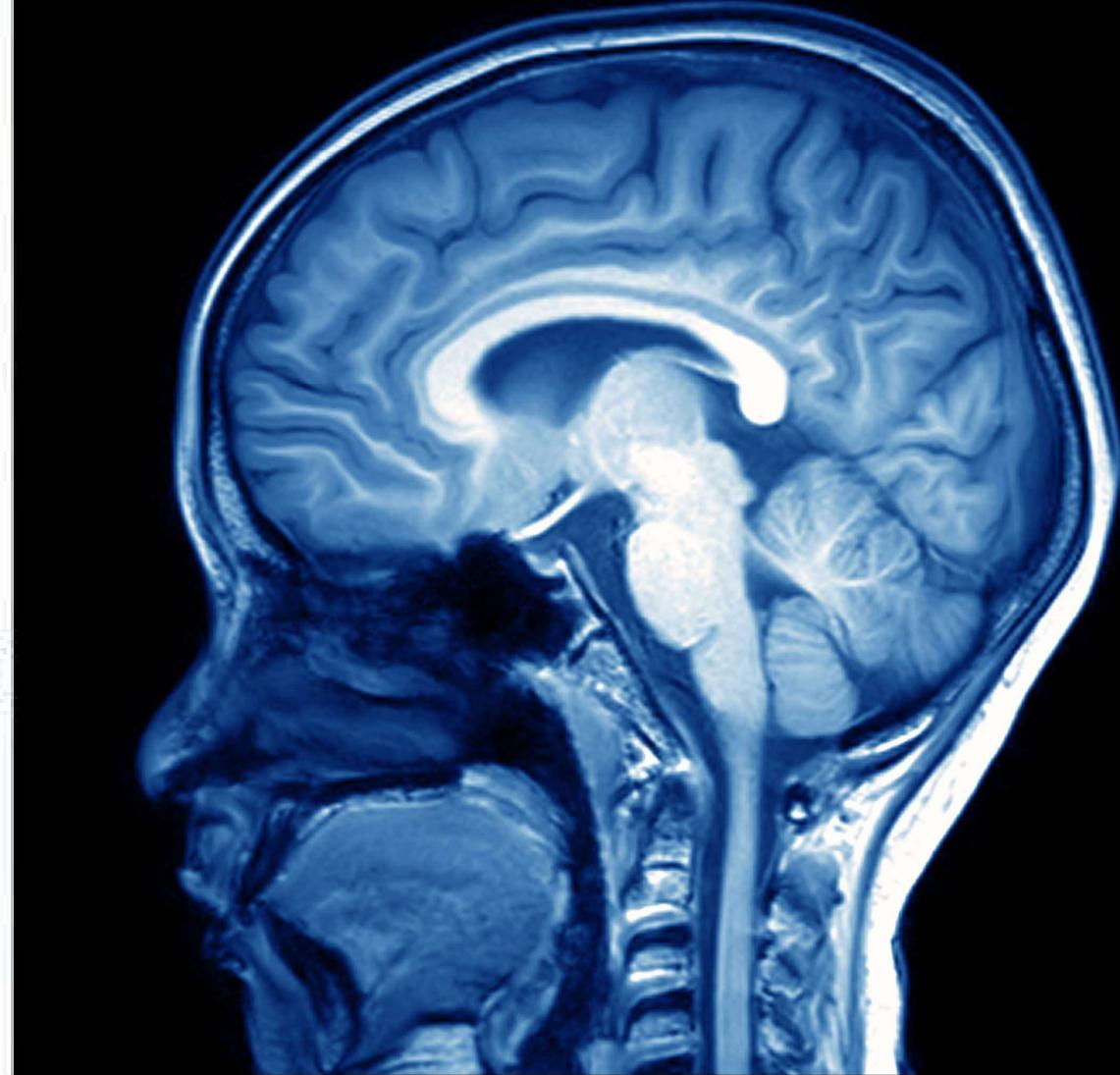
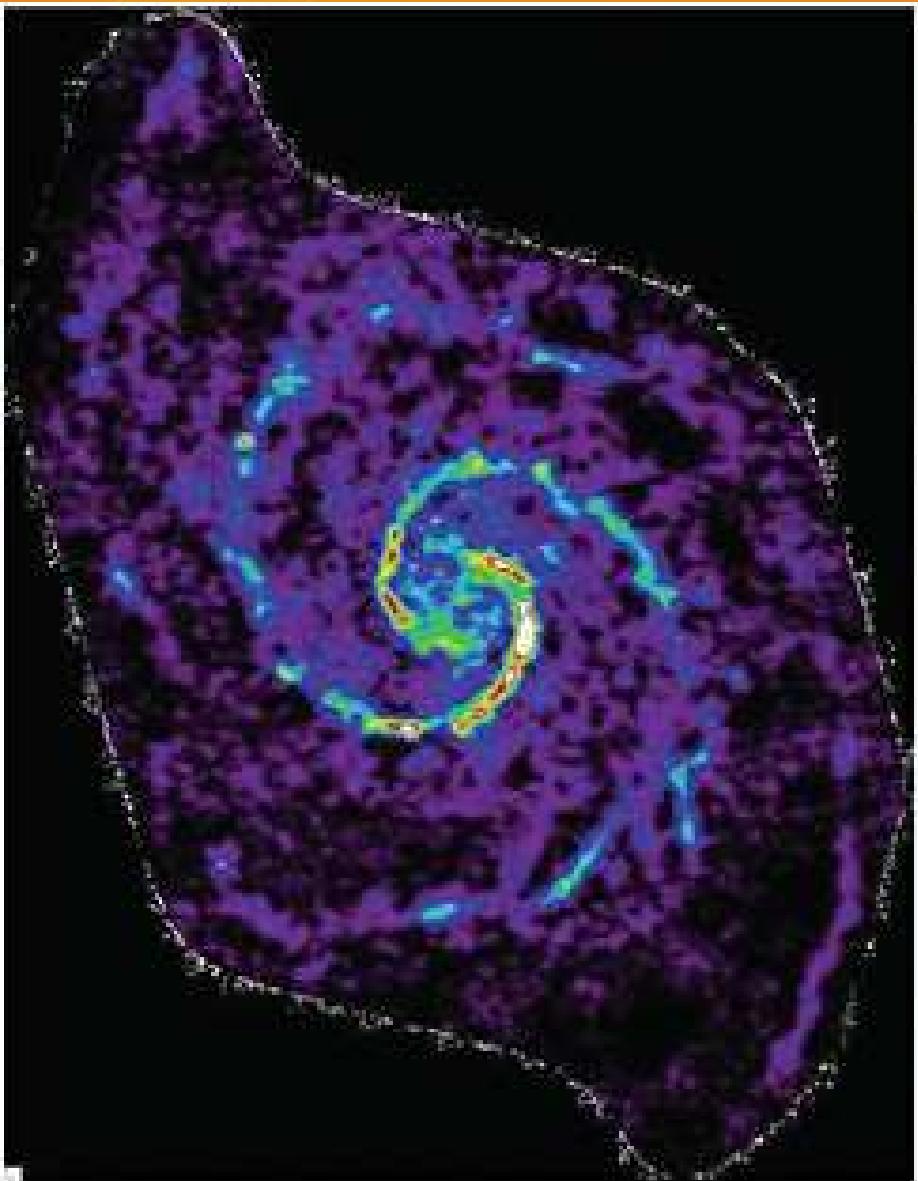
Time Keeping



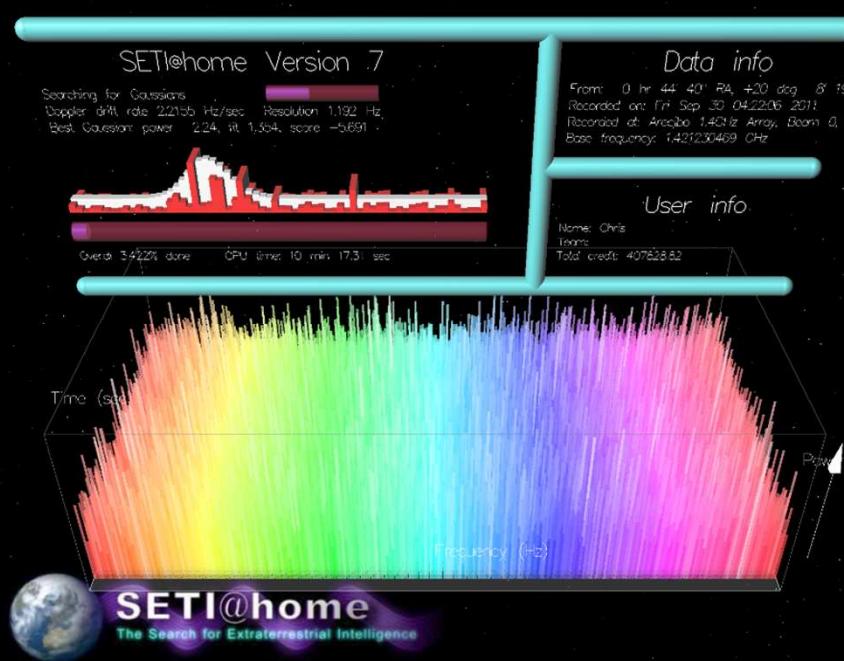
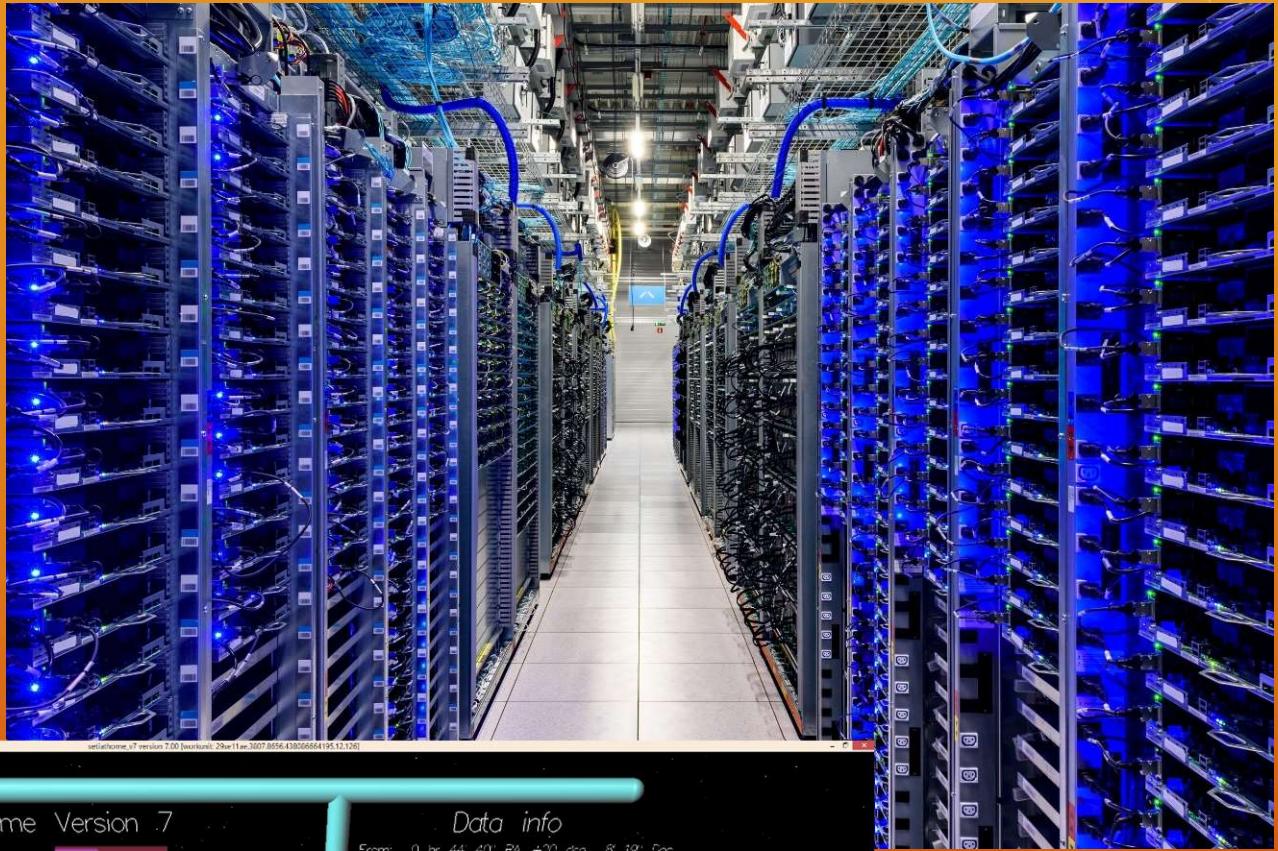
Amplifiers



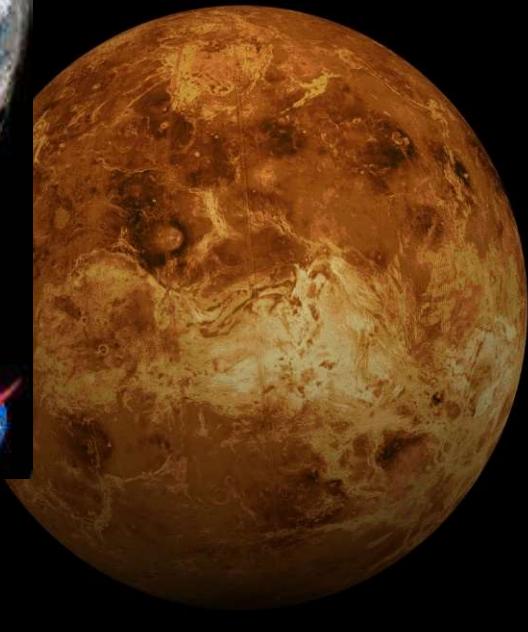
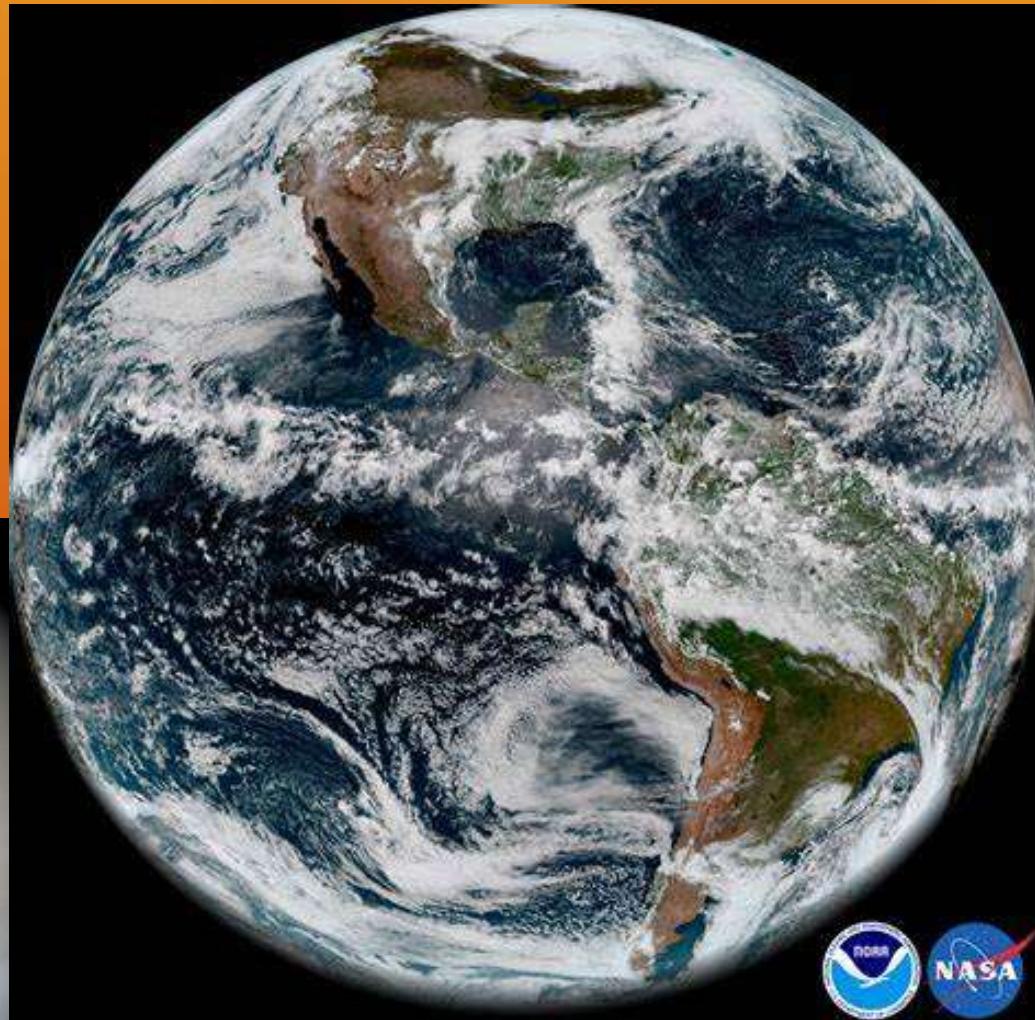
Imaging



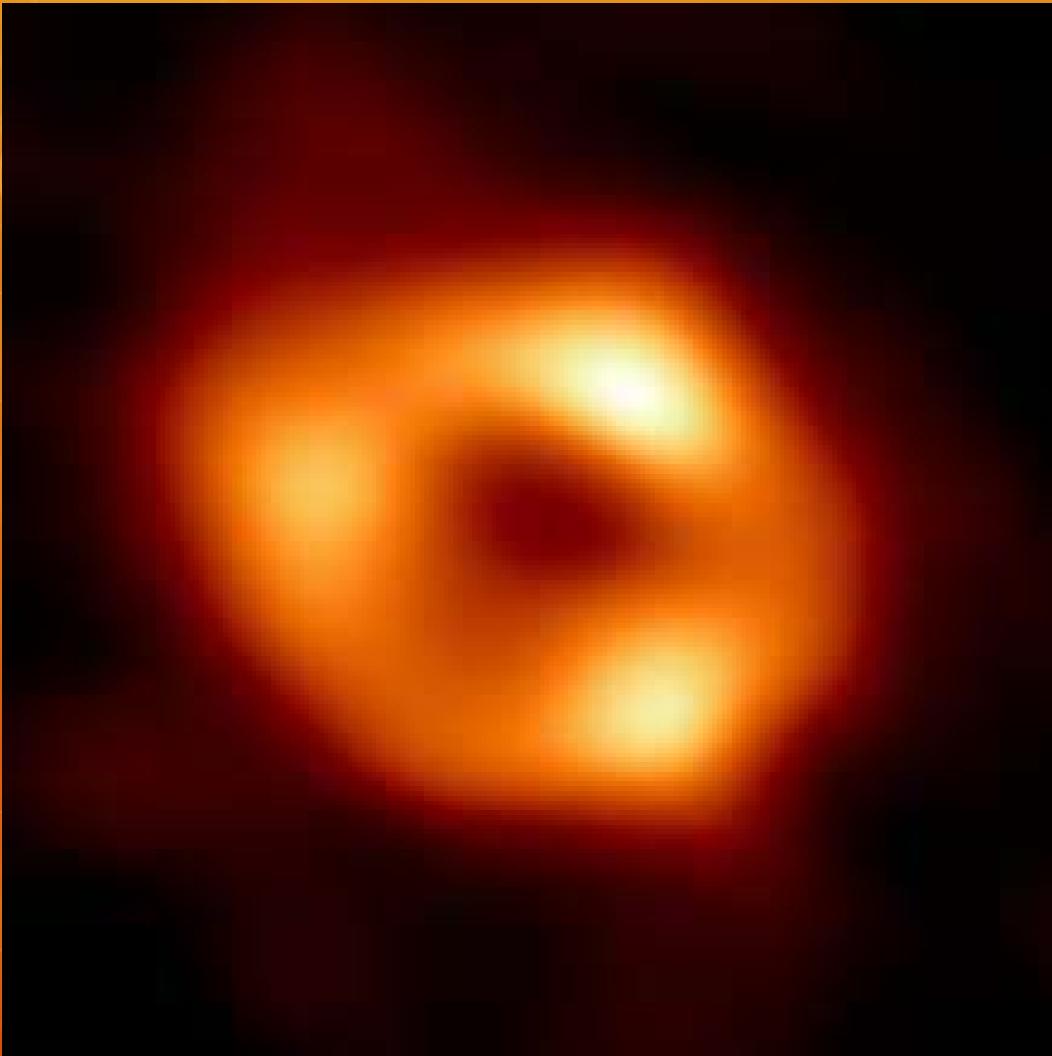
Computing



Astmospheric Change



Wi-Fi



Pure Curiosity





SARA Forum:
<https://groups.google.com/forum/#!forum/sara-list>



WVAS
<https://www.wvastro.org/>

Questions?